

PWA Formalism

M. Williams

Department of Physics
Carnegie Mellon University

February 10, 2006

Outline

1 Motivation

2 Integer Spin

- Spin-1 States at Rest
- Relativistic Spin-1 States
- Relativistic Spin-J States

3 Half-Integer Spin

- Relativistic Spin-1/2 States
- Relativistic Spin-J/2 States

4 Relativistic Orbital Angular Momentum

5 Rules for Writing PWA Amplitudes

6 Example Amplitude

- $\gamma p \rightarrow p \rho \rightarrow p \pi \pi$ t -channel

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Motivation

- Helicity formalism has awkward handling of final state particles with spin
- Helicity formalism is *NOT* relativistic
- We need to combine s -, t -, u -channel diagrams
- This formalism allows PWA parameters \rightarrow coupling constants
- C++ makes the code easy to deal with

General Technique

- Embed spin- j system, $(2j + 1)$ independent elements, into higher dimension space of rank- j tensors, 4^j independent elements.
- Supplementary conditions, Rarita-Schwinger, reduce independent elements to the proper number.
- Why do this?
 - Transformation properties of spin- j system become that of rank- j tensors.
 - Forming Lorentz scalars becomes straight forward.

Spin-1 States at Rest

Let $\vec{\epsilon}(\pm 1) = \mp \frac{1}{\sqrt{2}}(\hat{x} \pm i\hat{y})$ and $\vec{\epsilon}(0) = \hat{z}$.

Using the rotation operator about the k -axis, $e^{-i\theta J_k}$, and considering infinitesimal rotations we get $J_{ij}^k = -i\epsilon_{kij}$.

It then follows that,

- $J_z \vec{\epsilon}(m) = m \vec{\epsilon}$
- $J^2 \vec{\epsilon}(m) = 2 \vec{\epsilon}(m)$
- $\vec{\epsilon}^*(m) \cdot \vec{\epsilon}(m') = \delta_{mm'}$

Thus, $\vec{\epsilon}(m)$ are ortho-normal eigenstates of J^2 and J_z with $j = 1$ and $m_z = m$.

Also, note that $R\vec{\epsilon}(m) = \sum_{m'} D_{mm'}^{(1)}(R)\vec{\epsilon}(m')$.

Relativistic Spin-1 States

We now want to construct polarization 4-vectors (for massive particles).

- We need to reduce the number of independent components, the only measured quantity is p^μ , the constraint equation can be written as, $p^\mu \epsilon_\mu(p, m) = 0$.
- In the particle's rest frame, $p^\mu \epsilon_\mu = 0$ forces the energy component to zero. Thus, the rest frame states are $\epsilon_\mu(0, m) = (0, \vec{\epsilon}(m))$.
- To get the polarization 4-vector in any frame, simply boost it, $\epsilon^\mu(p, m) = \Lambda^\mu{}_\nu \epsilon^\nu(0, m)$

Relativistic Spin-1 States

- We can verify that $\epsilon^\mu(\mathbf{p}, m)$ has the correct rotational property,

$$\begin{aligned}
 R^\mu{}_\nu \epsilon^\nu(\mathbf{p}, m) &= R^\mu{}_\rho \Lambda^\rho{}_\pi(\mathbf{p}) \epsilon^\pi(\mathbf{p}_{rf}, m) \\
 &= R^\mu{}_\rho \Lambda^\rho{}_\pi(\mathbf{p}) (R^{-1})^\pi{}_\delta R^\delta{}_\sigma \epsilon^\sigma(\mathbf{p}_{rf}, m) \\
 &= \Lambda^\mu{}_\delta(R\mathbf{p}) \sum_{m'} D_{m'm}^{(1)}(R) \epsilon^\delta(\mathbf{p}_{rf}, m') \\
 &= \sum_{m'} D_{m'm}^{(1)}(R) \epsilon^\mu(R\mathbf{p}, m'),
 \end{aligned}$$

- We also define the spin-1 projection operator as,

$$P_{\mu\nu}^{(1)}(\mathbf{p}) = \sum_m \epsilon_\mu(\mathbf{p}, m) \epsilon_\nu^*(\mathbf{p}, m) = -g_{\mu\nu} + \frac{p_\mu p_\nu}{p^2} \equiv -\tilde{g}_{\mu\nu}.$$

Relativistic Spin-J States

The spin-J states are constructed out of the spin-1 states according to,

$$\epsilon_{\mu_1 \mu_2 \dots \mu_J}(\mathbf{p}, m) = \sum_{m_{J-1}, m_1} ((J-1)m_{J-1} 1 m_1 | J m) \\ \times \epsilon_{\mu_1 \mu_2 \dots \mu_{J-1}}(\mathbf{p}, m_{J-1}) \epsilon_{\mu_J}(\mathbf{p}, m_1).$$

- Rarita-Schwinger conditions for spin-J reduce the number of independent elements from 4^J to $(2J+1)$,

$$\begin{aligned} p^{\mu_i} \epsilon_{\mu_1 \mu_2 \dots \mu_i \dots \mu_J}(\mathbf{p}, m) &= 0 \\ \epsilon_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_J}(\mathbf{p}, m) &= \epsilon_{\mu_1 \mu_2 \dots \mu_j \dots \mu_i \dots \mu_J}(\mathbf{p}, m) \\ g^{\mu_i \mu_j} \epsilon_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_J}(\mathbf{p}, m) &= 0 \end{aligned}$$

- The spin-J projection operator is,

$$P_{\mu_1 \mu_2 \dots \mu_J \nu_1 \nu_2 \dots \nu_J}^{(J)}(\mathbf{p}) = \sum_m \epsilon_{\mu_1 \mu_2 \dots \mu_J}(\mathbf{p}, m) \epsilon_{\nu_1 \nu_2 \dots \nu_J}^*(\mathbf{p}, m)$$

General Technique

- Construct the relativistic spin-1/2 states using Dirac formalism
- Couple these to the spin-J states from the previous section to build higher half-integral spin states.

Relativistic Spin-1/2 States

Dirac formalism uses a 4-component spinor, $u(p, m)$, for spin-1/2 particles.

- The Dirac equation, $(\gamma^\mu p_\mu - w)u(p, m) = 0$, reduces the number of independent elements to 2.
- In the particle's rest frame, the spinor is $u(0, m) = \begin{pmatrix} \chi(m) \\ 0 \end{pmatrix}$.
- To get the spinor in any frame, boost it,

$$u(p, m) = \Lambda_{\frac{1}{2}}(p)u(0, m) = \sqrt{\frac{E+w}{2w}} \begin{pmatrix} \chi(m) \\ \frac{\vec{\sigma} \cdot \vec{p}}{E+w} \chi(m) \end{pmatrix},$$

- The spin-1/2 projection operator is, $P^{(\frac{1}{2})}(p) = \sum_m u(p, m)\bar{u}(p, m) = \frac{1}{2w}(\gamma^\mu p_\mu + w)$, where $\bar{u} \equiv u^\dagger \gamma^0$.

Useful Transformation Properties

Useful properties,

- $u \rightarrow \Lambda_{\frac{1}{2}} u$
- $\bar{u} \rightarrow \bar{u} \Lambda_{\frac{1}{2}}^{-1}$
- $\Lambda_{\frac{1}{2}}^{-1} \gamma^\mu \Lambda_{\frac{1}{2}} = \Lambda^\mu{}_\nu \gamma^\nu$

Transformation properties of $\bar{u} \Gamma u$,

- $\Gamma = 1$ *scalar*
- $\Gamma = \gamma^5$ *pseudo-scalar*
- $\Gamma = \gamma^\mu$ *vector*
- $\Gamma = \gamma^\mu \gamma^5$ *pseudo-vector*
- $\Gamma = \gamma^\mu \gamma^\nu$ *tensor*

Relativistic Spin-J/2 States

The spin-J/2 states are constructed out of the spin-1/2 and spin-n, where $J/2 = n + 1/2$, states,

$$u_{\mu_1 \mu_2 \dots \mu_n}(\mathbf{p}, m) = \sum_{m_n m_{\frac{1}{2}}} (n m_n \frac{1}{2} m_{\frac{1}{2}} | J m) \epsilon_{\mu_1 \mu_2 \dots \mu_n}(\mathbf{p}, m_n) u(\mathbf{p}, m_{\frac{1}{2}}),$$

- The Rarita-Schwinger conditions for spin-J/2 are,

$$(\gamma^\mu p_\mu - w) u_{\mu_1 \mu_2 \dots \mu_n}(\mathbf{p}, m) = 0$$

$$u_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_n}(\mathbf{p}, m) = u_{\mu_1 \mu_2 \dots \mu_j \dots \mu_i \dots \mu_n}(\mathbf{p}, m)$$

$$p^{\mu_i} u_{\mu_1 \mu_2 \dots \mu_i \dots \mu_n}(\mathbf{p}, m) = 0$$

$$\gamma^{\mu_i} u_{\mu_1 \mu_2 \dots \mu_i \dots \mu_n}(\mathbf{p}, m) = 0$$

$$g^{\mu_i \mu_j} u_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_n}(\mathbf{p}, m) = 0$$

- The projection operator is,

$$P_{\mu_1 \mu_2 \dots \mu_n \nu_1 \nu_2 \dots \nu_n}^{(J/2)}(\mathbf{p}) = \sum_m u_{\mu_1 \mu_2 \dots \mu_n}(\mathbf{p}, m) \bar{u}_{\nu_1 \nu_2 \dots \nu_n}(\mathbf{p}, m)$$

Relativistic Orbital Angular Momentum

Consider $X \rightarrow a + b$, with daughter momenta p_a, p_b , total momentum $P = p_a + p_b$ and relative momentum $p_{ab} = \frac{1}{2}(p_a - p_b)$.

- Construct a state of pure angular momentum, ℓ ,

$$L_{\mu_1 \mu_2 \dots \mu_\ell}^{(\ell)}(p_{ab}) = (-)^\ell P_{\mu_1 \mu_2 \dots \mu_\ell \nu_1 \nu_2 \dots \nu_\ell}^{(\ell)}(P) p_{ab}^{\nu_1} p_{ab}^{\nu_2} \dots p_{ab}^{\nu_\ell}.$$

- $L_{\mu_1 \mu_2 \dots \mu_\ell}^{(\ell)}$ satisfies the Rarita-Schwinger conditions (it has $(2\ell + 1)$ independent elements),

$$P^{\mu_i} L_{\mu_1 \mu_2 \dots \mu_i \dots \mu_\ell}^{(\ell)}(p_{ab}) = 0$$

$$L_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_\ell}^{(\ell)}(p_{ab}) = L_{\mu_1 \mu_2 \dots \mu_j \dots \mu_i \dots \mu_\ell}^{(\ell)}(p_{ab})$$

$$g^{\mu_i \mu_j} L_{\mu_1 \mu_2 \dots \mu_i \dots \mu_j \dots \mu_\ell}^{(\ell)}(p_{ab}) = 0,$$

Orbital Tensor Example

Let's look at this for $\ell = 2$.

- $L_{\mu_1\mu_2}^{(2)}(p_{ab}) = \tilde{p}_{\mu_1}^{ab}\tilde{p}_{\mu_2}^{ab} - \frac{1}{3}\tilde{p}_{ab}^2\tilde{g}_{\mu_1\mu_2}$, where $\tilde{g}_{\mu_1\mu_2} = g_{\mu_1\mu_2} - \frac{P_{\mu_1}P_{\mu_2}}{P^2}$
and $\tilde{p}_{\mu}^{ab} = \tilde{g}_{\mu\nu}p_{ab}^{\nu}$.
- In the CM frame, $\tilde{p}_{\mu}^{ab} \rightarrow \vec{p}_a = \vec{p}$ and $L_{\mu_1\mu_2}^{(2)} \rightarrow L_{ij}^{(2)} = p_i p_j - \frac{1}{3}p^2 \delta_{ij}$.
- Project out an M_z piece,

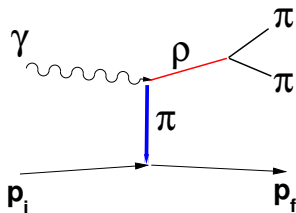
$$\begin{aligned}
 \langle L = 2 | M_z = 1 \rangle &\rightarrow \langle L = 2 | m_1 = 1; m_2 = 0 \rangle \\
 &\rightarrow L_{ij}^{(2)} \epsilon^i(1) \epsilon^j(0) \\
 &= (\vec{p} \cdot \vec{\epsilon}(1))(\vec{p} \cdot \vec{\epsilon}(0)) - \frac{1}{3}p^2 \vec{\epsilon}(1) \cdot \vec{\epsilon}(0) \\
 &= p^2 \cos(\theta) \sin(\theta) e^{i\phi} \\
 &\sim p^2 Y_{21}(\theta, \phi)
 \end{aligned}$$

Rules for Writing PWA Amplitudes

- Amplitude must be a Lorentz Scalar, gauge invariant (current conserving) and, where applicable, conserve parity.
- initial (final) spin-0 particles $\rightarrow 1$
- initial (final) spin-1/2 particles $\rightarrow u(p, m) (\bar{u}(p, m))$
- initial (final) spin-1 particles $\rightarrow \epsilon^\mu(p, m) (\epsilon^{*\mu}(p, m))$
- intermediate spin-j particles $\rightarrow P^{(j)}(p)$
- If a pure orbital angular momentum amplitude is needed, then *all* factors of p_a, p_b must in $L_{\mu_1 \mu_2 \dots \mu_\ell}^{(\ell)}(p_{ab})$.
- The only other allowed components are $g_{\mu\nu}, \epsilon_{\mu\nu\rho\sigma}, \gamma^\mu, \gamma^5$, factors of momentum (if not looking at a specific orbital state) and any Lorentz scalars (complex constants, p^2, \dots).

$\gamma\rho \rightarrow p\rho \rightarrow p\pi\pi$ t -channel

A simple example that could apply to Gluex, ρ photo-production off the proton. In this example, we'll consider only π exchange in t -channel.



- First consider $\rho \rightarrow \pi\pi$, which proceeds only with $\ell = 1$. The amplitude can be written as,

$$L_{\mu}^{(1)}(p_{\pi\pi})\epsilon^{\mu}(p_{\rho}, m_{\rho})$$

$\gamma\rho \rightarrow \rho\rho \rightarrow \rho\pi\pi$ t -channel

- The $\gamma\pi\rho$ vertex must look like,

$$\epsilon^{*\mu}(\mathbf{p}_\rho, m_\rho) X_{\mu\nu} \epsilon^\nu(\mathbf{p}_\gamma, m_\gamma)$$

Parity conservation $\rightarrow X_{\mu\nu}$ has *negative* parity, so we write it as,

$$X_{\mu\nu} = \epsilon_{\mu\nu\alpha\beta} p_\rho^\alpha p_\gamma^\beta$$

Therefore, the amplitude is written as,

$$\epsilon^{*\mu}(\mathbf{p}_\rho, m_\rho) \epsilon_{\mu\nu\alpha\beta} p_\rho^\alpha p_\gamma^\beta \epsilon^\nu(\mathbf{p}_\gamma, m_\gamma)$$

$\gamma p \rightarrow p \rho \rightarrow p \pi \pi$ t -channel

- The $pp\pi$ vertex must look like,

$$\bar{u}(p_f, m_f) X u(p_i, m_i)$$

Parity conservation \rightarrow X has *negative* parity, so we write the amplitude as,

$$\bar{u}(p_f, m_f) \gamma^5 u(p_i, m_i)$$

$\gamma\rho \rightarrow p\rho \rightarrow p\pi\pi$ t -channel

- Putting this all together, the amplitude is,

$$L_{\sigma}^{(1)}(p_{\pi\pi}) \left(\sum_{m_{\rho}} \epsilon^{\sigma}(p_{\rho}, m_{\rho}) \epsilon^{*\mu}(p_{\rho}, m_{\rho}) \right) \epsilon_{\mu\nu\alpha\beta} p_{\rho}^{\alpha} p_{\gamma}^{\beta} \epsilon^{\nu}(p_{\gamma}, m_{\gamma}) \\ \times \bar{u}(p_f, m_f) \gamma^5 u(p_i, m_i) BW(p_{\rho}) R_{\pi}(p_t)$$

- Which can be written using the spin-1 projection operator as,

$$L_{\sigma}^{(1)}(p_{\pi\pi}) P^{(1)\sigma\mu}(p_{\rho}) \epsilon_{\mu\nu\alpha\beta} p_{\rho}^{\alpha} p_{\gamma}^{\beta} \epsilon^{\nu}(p_{\gamma}, m_{\gamma}) \\ \times \bar{u}(p_f, m_f) \gamma^5 u(p_i, m_i) BW(p_{\rho}) R_{\pi}(p_t)$$

$\gamma p \rightarrow p \rho \rightarrow p \pi \pi$ t -channel Code

How do we code up this example?

- The amplitude,

$$L_{\sigma}^{(1)}(p_{\pi\pi}) P^{(1)\sigma\mu}(p_{\rho}) \epsilon_{\mu\nu\alpha\beta} p_{\rho}^{\alpha} p_{\gamma}^{\beta} \epsilon^{\nu}(p_{\gamma}, m_{\gamma}) \\ \times \bar{u}(p_f, m_f) \gamma^5 u(p_i, m_i) BW(p_{\rho}) R_{\pi}(p_t)$$

- The code,

```
L1*rho.Projector()*(levi_civita|(p4_rho%p4_gamma%eps(m_gamma)))
Bar(uf(m_f))*gamma5*ui(m_i)*BrightWigner(p4_rho,mass_rho,width_rho)
*ReggePropagator(t,s,0,pi_slope,pi_inter,pi_sig)
```

Summary

- Tensor formalism is Relativistic
- Tensor formalism handles s -, t -, u -channel in a natural way
- PWA parameters \rightarrow physics couplings
- We've developed code to take care of the difficulties...so why not use it?